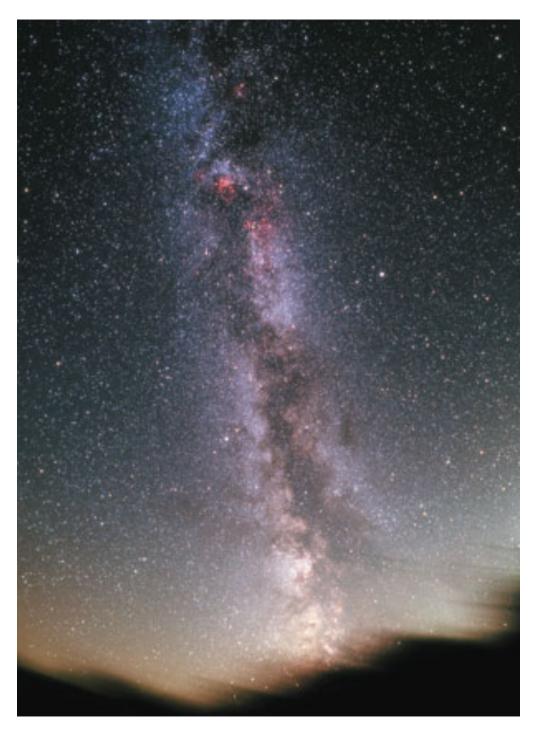
# The Milky Way

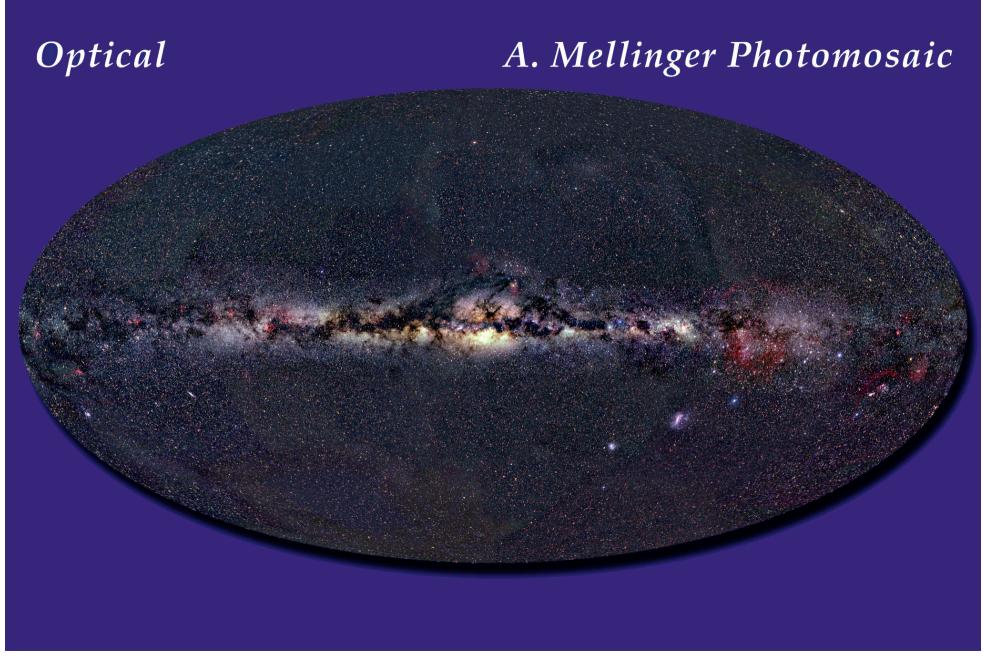
- The Galactic Plane
- Spiral arms
- Galactic Halo

### The Galactic Plane

- In optical light the Milky Way delineates the plane of our Galaxy
- Dust lanes obscure much of the structure
- The true structure is only revealed in the near-IR where extinction is lower
- We are located about 8 kpc from the centre in a disc with a radius of about 20 kpc

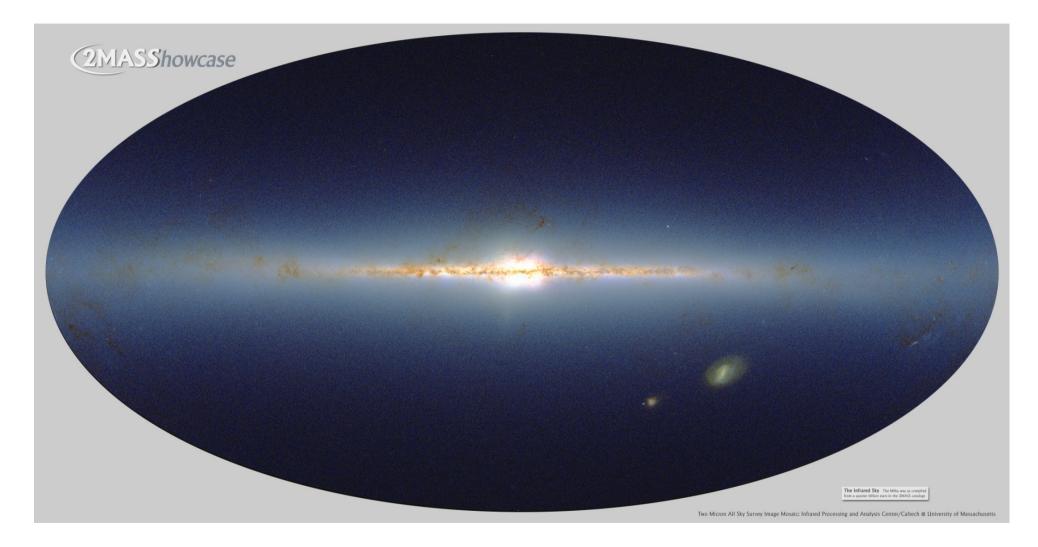


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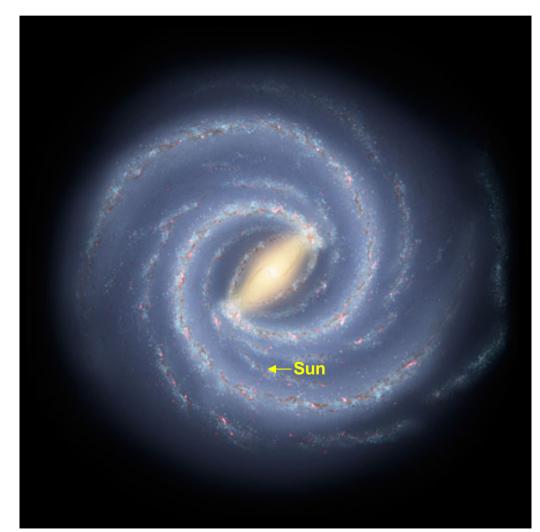
#### Near-infrared: 1-2 $\mu$ m



Atlas Image obtained as part of the Two Micron All Sky Survey (2MASS), a joint project of the University of Massachusetts and the Infrared Processing and Analysis Center/California Institute of Technology, funded by the National Aeronautics and Space Administration and the National Science Foundation.

### The Bulge and Bar

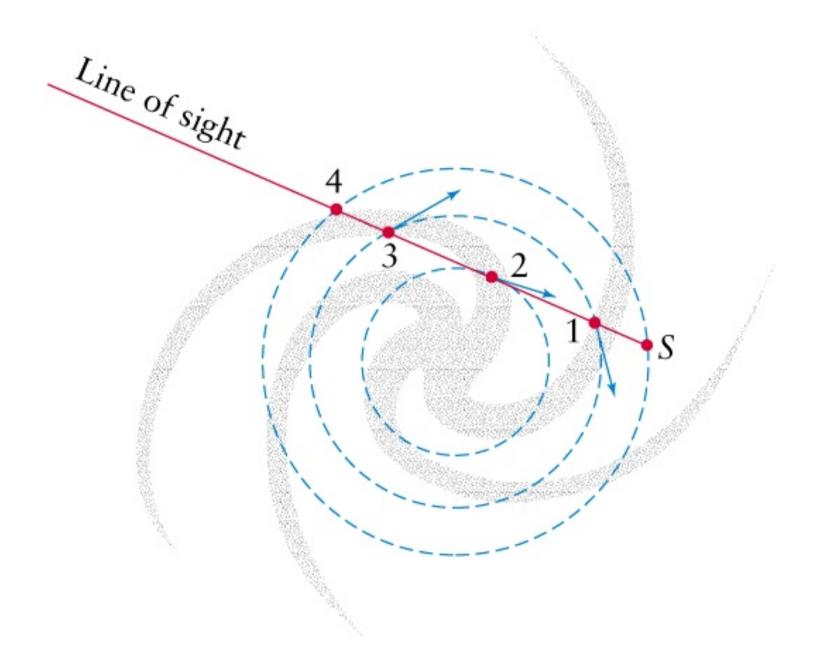
- Bulge is about 1 kpc in radius
- Bar is about 4 kpc in length

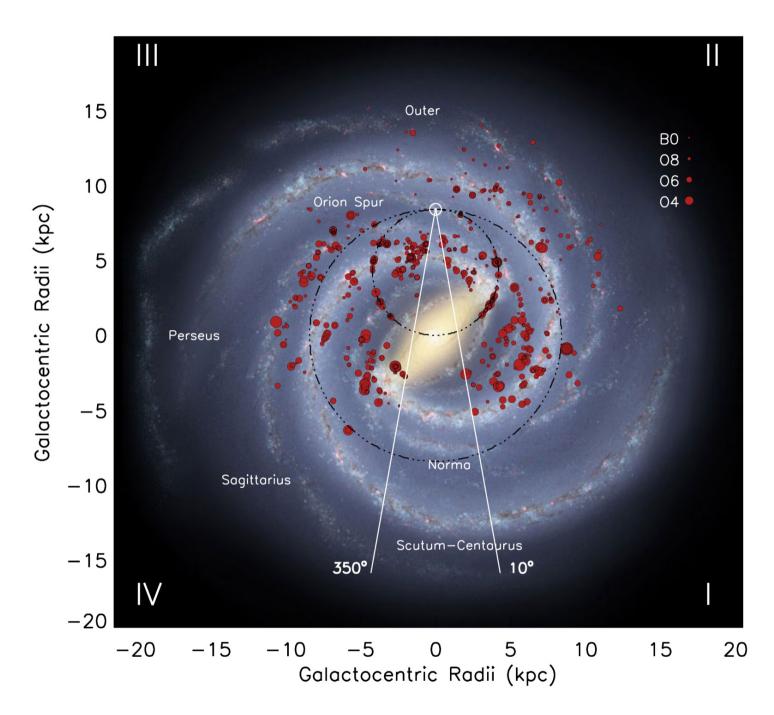


credit: NASA/JPL-Caltech

#### **Spiral Arms**

- The spiral arm structure is determined using tracers of massive star formation with known distances – usually via their Doppler shift and a rotation model
- H II regions and CO clouds are used
- Our galaxy has 4 star-forming gaseous arms and 2 stellar arms originating at each end of the bar





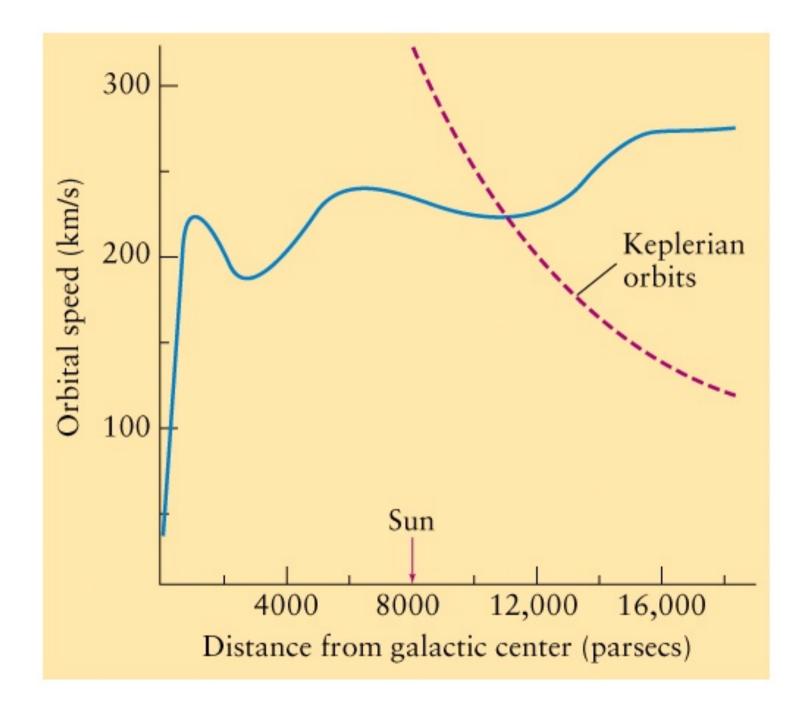
Distribution of massive star forming regions from the Red MSX Source Survey www.ast.leeds.ac.uk/RMS

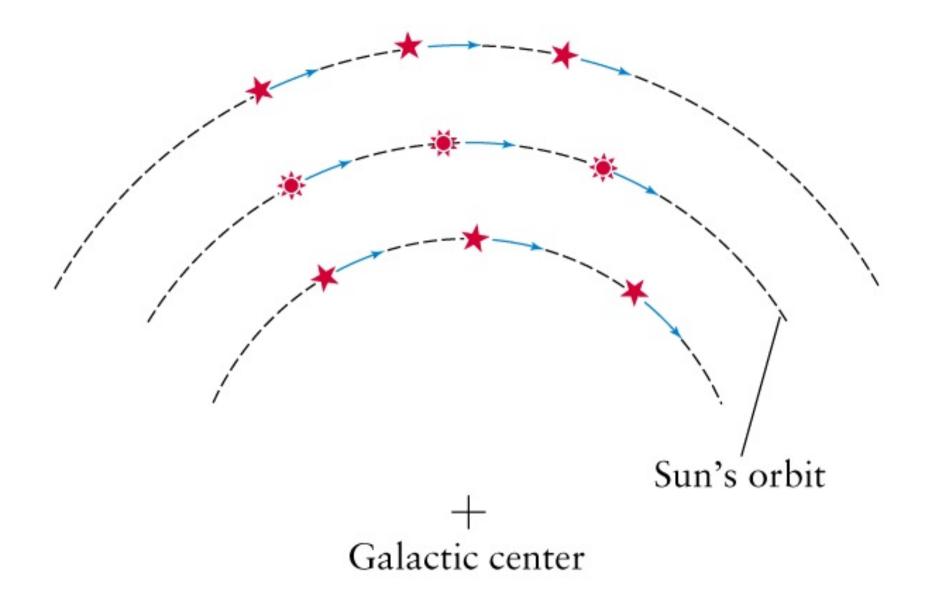
#### Rotation of the disc

 The observed rotational velocity of the Galactic disc is approximately constant with radius

 $v \approx \text{constant} \approx 220 \text{ km s}^{-1}$ 

• This requires differential rotation: stars closer to the centre of the galaxy orbit in a shorter time (period P  $\alpha$  r) and overtake us, whilst we overtake stars further out

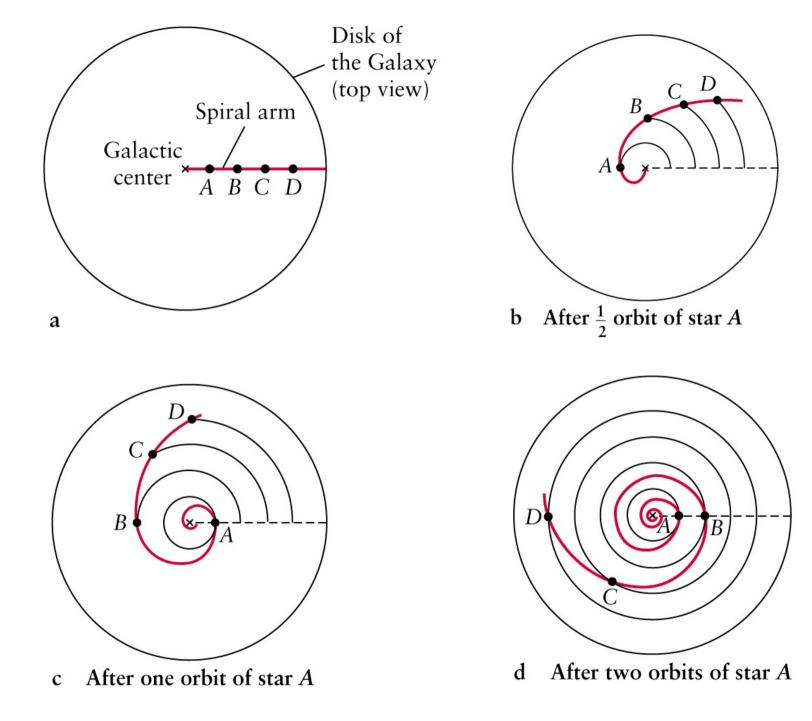




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#### The Wind-up Problem

 If the spiral arms were a fixed pattern in the stars and gas the differential rotation would cause them to 'wind-up' in a few revolutions



### **Density Waves**

- The current model for spiral arms is that they are a density wave pattern that rotates at a slower speed than the galaxy
- Stars and gas pass in and out of the arm
- As gas gets compressed in the arm molecular clouds form with subsequent star formation
- Spiral arms usually trail the rotation

 Dust lane where material enters spiral arm, then H II regions, then blue stars



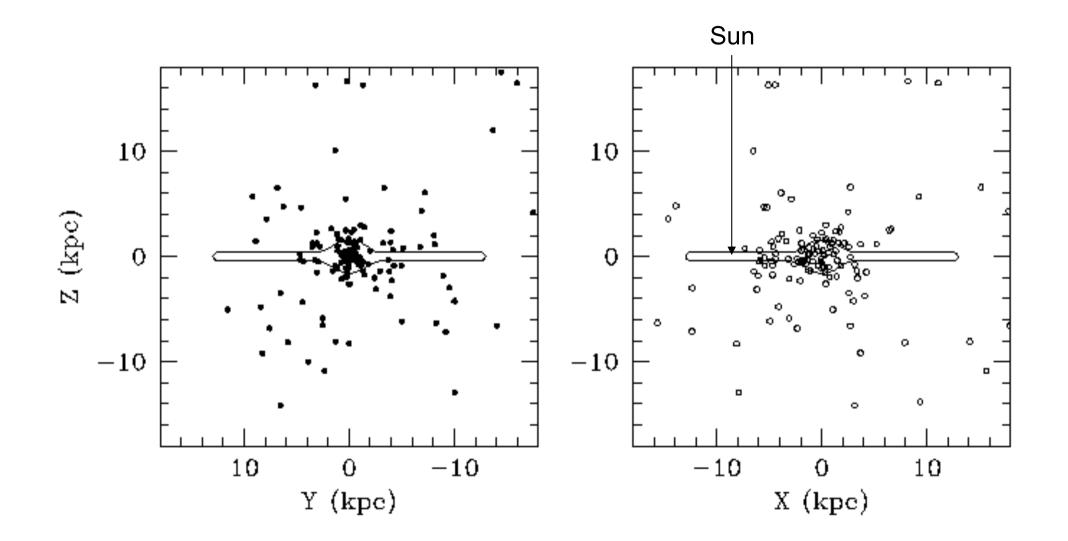
Credit: NASA and The Hubble Heritage Team (STScI/AURA)

#### The Galactic Halo

- The plane of our galaxy is surrounded by a more spherical halo of objects
- Consists of globular clusters and halo stars
- Total radius of the halo is ~ 100 kpc

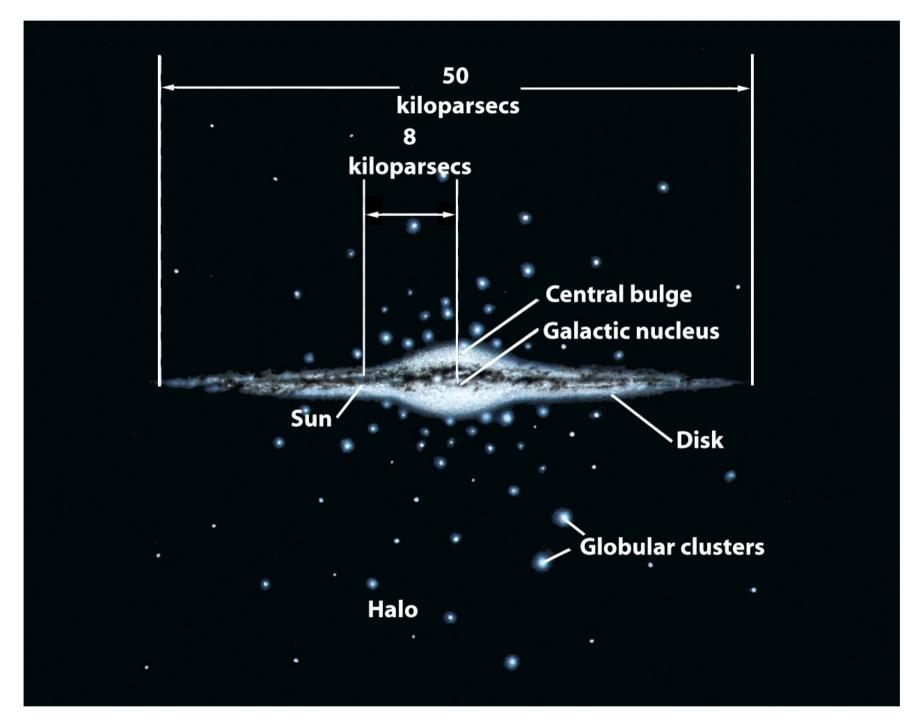


Globular Cluster M80: NASA HST



Distribution of nearby globular clusters in the plane of the Galaxy

Harris, W. E. "Globular Clusters", 1999, Cambridge University Press



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## Formation of the Galaxy

- Spherical halo formed first out of metalpoor material
  - One initial burst of star formation and none since
- Disc formed later
  - Continuous star formation leading to metalrich population
- Bulge also has some metal-rich stars as a result of mergers with small galaxies

#### Summary

- Our Galaxy is a SBb or c
- The disc is surrounded by a spherical galactic halo
- The galactic halo is made up of Population II objects and was formed at an earlier stage than the disc